Search for Baryon Number Violation in Neutron Oscillations

Leah Broussard Oak Ridge National Laboratory

XIV Latin American Symposium on Nuclear Physics and Applications (LASNPA 2024) June 17-21, 2024





What makes up 85% of matter in the universe? Why is there no antimatter in the universe?





Symmetry tests, rare processes, and precision

measurements



Leah Broussard

"Search for Baryon Number Violation in Neutron Oscillations," LASNPA 2024, June 17 – 21, 2024

Baryon and Lepton Number



Quarks: $\mathbf{B} = +\frac{1}{3}$ Antiquarks: $\mathbf{B} = -\frac{1}{3}$



Electrons/ Neutrinos: L = +1Anti-electrons/ Anti-neutrinos: L = -1



"We must have a process that does not conserve **B** [or **L**] to create an excess of matter in the universe!"

Andrei Sakharov

...and we might expect these to be violated! B - L conserved in the SM $\Delta B \neq 0, \Delta L = 0, \Delta [B - L] \neq 0$ $\Delta B = 0, \Delta L \neq 0, \Delta [B - L] \neq 0$ $\Delta B \neq 0, \Delta L \neq 0, \Delta [B - L] = 0$

Searches for Violation of B and L

- **Proton Decay**: Originally motivated massive detectors!
 - But limit is now $> 10^{34}$ years
 - Universe is 10¹⁰ years old

$$0\upsilon 2\beta$$

$$\Delta B = 0, \Delta L \neq 0$$

$$\Delta [B - L] \neq 0$$

 $p \rightarrow e^{+} + \pi^{0}$ $\Delta B \neq 0, \Delta L \neq 0$

 $\Delta[\alpha - 3] = 0$

$$n \to \overline{n}$$

$$\Delta B = 2, \Delta L = 0$$

$$\Delta [B - L] \neq 0$$







Leah Broussard

Neutron oscillations $n \rightarrow \overline{n}$

• Two-state mixing

$$\widehat{H}\psi = \begin{pmatrix} H_n & \epsilon \\ \epsilon & H_{\overline{n}} \end{pmatrix} \begin{pmatrix} n \\ \overline{n} \end{pmatrix} \qquad H_n = m_n + \frac{p^2}{2m_n} + \mu_n(\overrightarrow{\sigma} \cdot \overrightarrow{B}) + V - i W - \frac{i}{2\tau_n}$$
Antiparticles have opposite magnetic moments Different material potentials

- Oscillations are suppressed unless energy difference $\ll \epsilon < 10^{-29}\,{\rm MeV}$
- Use uncertainty principle $\Delta E \Delta t \ll \hbar$ to operate in "quasi-free limit" <u>NIMA 320 (1992) 569</u>
- Free neutrons
 - Earth magnetic potential $\mu \cdot B$ (10⁻¹⁸ MeV)
 - Free neutron TOF ~ 0.1 s \rightarrow B~10 nT
 - $P_{n \to \overline{n}}(t) = \left(\frac{t_{free}}{\tau_{n \to \overline{n}}}\right)^2$
- Free neutron Hamiltonian

Signature of $n \to \overline{n}$



- Best published limits from **Super-K**
 - 11 candidate events, 9.3 expected bkgds







• DUNE future reach: $\tau_{n \to \overline{n}} > 5.53 \times 10^8$ s <u>arXiv:2002.03005</u>

Leah Broussard

The NNBAR Experiment





Recent NNBAR progress



Complete engineering design of ESS second cold neutron source, optimized for fundamental physics HighNESS CDR Vol 1: arXiv:2309.17333



Proposed sensitivity can be achieved even with ESS at 2 MW



NNBAR CDR completed with conceptual engineering Detector an design: beamline, reflector, magnetic and radiation shielding, detector (JNR **25** 315-406 (2024))

Unique "Large Beam Port" is already constructed, earmarked for NNBAR

The NNBAR experiment at the **European Spallation Source**

beam stop

Leah Broussard

"Search for Baryon Number Violation in Neutron Oscillations," LASNPA 2024, June 17 – 21, 2024

eamlin

Antineutrons and dark neutrons?

- NNBAR = well motivated large-scale O(100M) project planned for >2030
- Complementary science on a shorter timescale: $\Delta B = 1$ neutron oscillations
- Straightforward extension of formalism: consider $n \rightarrow \overline{n}$, n', \overline{n}'
- Connection to cobaryogenesis IJMP **33** 1844034 (2018)
- "Shortcut" for neutronantineutron oscillations via dark sector <u>EPJC **81** (2021) 33</u>

$$\mathcal{H}_{int} = \begin{pmatrix} m + \mu \boldsymbol{\sigma} \cdot \boldsymbol{B} & \varepsilon_{n\overline{n}} & \alpha_{nn'} & \delta_{n\overline{n}'} \\ \varepsilon_{n\overline{n}} & m - \mu \boldsymbol{\sigma} \cdot \boldsymbol{B} & \delta_{n\overline{n}'} & \alpha_{nn'} \\ \alpha_{nn'} & \delta_{n\overline{n}'} & m' + \mu' \boldsymbol{\sigma} \cdot \boldsymbol{B}' & \varepsilon_{n\overline{n}} \\ \delta_{n\overline{n}'} & \alpha_{nn'} & \varepsilon_{n\overline{n}} & m' - \mu' \boldsymbol{\sigma} \cdot \boldsymbol{B}' \end{pmatrix}$$

• Opportunity for NNBAR R&D needed now in detector prototyping, data selection algorithms, interface of beamline and detector

 $|\phi_n\rangle =$

Staged Program to NNBAR

- Increasingly complex searches:
 - 1. $n \rightarrow n' \rightarrow n$ with different assumptions:
 - A. Mass splitting between n' and n
 - B. "Transition magnetic moment" enhances probability
 - C. "Traditional" search for $n \rightarrow n' \rightarrow n$
 - *2.* $n \rightarrow n' \rightarrow \overline{n}$ in HIBEAM
 - *3.* $n \rightarrow \overline{n}$ in NNBAR
- Cold neutron regeneration technique demonstrated at ORNL

Perform at ORNL using existing neutron scattering instruments

Perform at ESS using dedicated instrument for high sensitivity



Regenerated neutrons (or antineutrons) are detectable

Mirror matter and mirror neutrons

- Decades of searches for dark matter: no confirmed detection. New avenues needed!
- Mirror Matter: identical particles/forces, opposite parity [Phys.Usp. 50 (2007) 380-389, From Fields to Strings 3 (2015) 2147, Phys.Rev. 104 (1956) 254-258]
- Only interacts with known particles via gravity [Sov.J.Nucl.Phys. 3 (1966) 6]
- A "hidden sector" which could help explain dark matter [PLB 503 (2001) 362, IJMPA 29 (2014) 1430013]
- Experimental observable: predictions of neutron oscillations in Mirror Matter models <u>PRL 96 081801 (2006)</u>; related to cobaryogenesis <u>IJMP 33 1844034 (2018)</u>
- Neutron could be one of a few portals to a dark sector!



Neutron oscillations $n \rightarrow n'$

• Two state mixing



- As before, as $\Delta E \rightarrow 0$, probability for oscillation resonantly enhanced
- Rare process, but not background-free. Tune laboratory B or V to look for evidence of Δm , B', $\epsilon_{nn'}$...

Prior/ongoing searches for mirror neutrons

- Strong limits from B dependence of stored ultracold neutrons
 - $\tau_{nn'} > 448 \text{ s} (90\% \text{ CL}) \text{ assuming } \Delta E = 0 \text{ (e.g. no mirror field } B') \underline{NIMA 611 (2008) 137}$



Can n - n' explain n lifetime anomaly?

- "Bottle" technique:: use a bottle of ultracold neutrons, observe surviving neutrons
- "Beam" technique: use a beam of cold neutrons, count decays into protons
- 4σ discrepancy persists between "beam"⁸⁸⁰ and "bottle" measurements <u>Atoms 6 (2018)</u> 875 <u>4</u>
- Do neutrons disappear into something other than protons?



$n \rightarrow n'$ in the NIST Beam Lifetime experiment



Leah Broussard

$n \rightarrow n'$ in the NIST Beam Lifetime experiment

 $\widehat{H} = \begin{pmatrix} \Delta m + V - iW \pm \mu B & \epsilon \\ \epsilon & 0 \end{pmatrix}$

 $\hat{\rho}(0) = \begin{pmatrix} 1 & 0 \\ 0 & 0 \end{pmatrix} = n$

• Calculated evolution through NIST beam lifetime B profile <u>PRC 71 (2005) 055502</u>

 $\frac{\partial}{\partial t}\hat{\rho} = -i[\hat{H}\cdot\hat{\rho}] = -i\hat{H}\hat{\rho} + i\hat{\rho}\hat{H}^{\dagger}$ Liouville-von Neumann equation

• Combine calculated shift in \dot{N}_p , \dot{N}_n to find a lifetime shift

 $\tau_{meas} = \frac{L}{v_n} \frac{N_n/\epsilon_n}{N_n/\epsilon_n}$

- Region of interest for $n \rightarrow n'$:
 - When $\delta \tau_{meas}/\tau_n \sim 1\%$



$n \rightarrow n' \rightarrow n@$ SNS experiment approach



Leah Broussard

"Search for Baryon Number Violation in Neutron Oscillations," LASNPA 2024, June 17 – 21, 2024

Impact for neutron lifetime anomaly

B Field at B4C	B4C Position	Total Cts in ROI	Charge	Cts / 5C	10 1			For $\Delta m \gg \mu B$		
+4.8 T	center	7748 ± 88	13.8 C	2805 ± 32	10-2			probab	$ility \sim cons$	
-4.8 T	center	4976 ± 70	8.8 C	2823 ± 4 0	10		1			
0 T	center	6631 ± 81	11.9 C	2791 ± 34	-	The second se	1	//		
3.33 T	center	1017 ± 32	1.8 C	2817 ± 88	€ 10 ⁻³		1	1		
+6.6 T	peak B	1010 ± 32	1.8 C	2804 ± 88		N N	No.		95% CL	
-6.6 T	peak B	1120 ± 33	2.0 C	2863 ± 86	10-4	- r	1		$ P = 10^{-5}$	
0 T	peak B	4916 ± 70	8.7 C	2815 ± 4 0)		$ P = 10^{-8}$	
• Exclu	ded trar	nsmission	n < 2.5	x 10 ⁻⁸ (95	⁰⁄0 ₁₀ -5 ⊥	F	<u>}</u>		UCN Tra	

- C.L.) (grey region)
- **Conclusion**: this exotic process does not explain neutron lifetime anomaly PRL 128 (2022) 212503

Gray – Excluded region from our experiment (95% CL) Red – 1% \pm 0.2% difference in neutron lifetime Dashed – Probability bands

 Δm (neV)

F. Gonzalez, ORNL

Improved search for $n \rightarrow n'(a)$ SNS

- ×10 reduction in detector background
- Higher intensity (better neutron beam acceptance) and improved analysis
- Improved understanding of neutron oscillations in matter → better choice of material for full sensitivity



Cadmium neutron absorber





Final result for $n \rightarrow n'$ @ SNS



Next: searches for $n \rightarrow n'$ (*a*) HFIR

- High Flux Isotope Reactor 85 MW: highest reactor-based source of neutrons for research in US
 - ~10,000× more neutron intensity than SNS
 - Lower backgrounds: ³He neutron detector in Cd shielded tank
 - 15 m and 20 m long beamguides for "disappearance" and "regeneration"





$n \rightarrow n'$ (*a*) HFIR program now launched

- SNS experiment setup moved to HFIR and completed data-taking in January 2024
 - Sensitivity gain primarily due to higher neutron flux
 - Stage 1A complete!
- Now prototyping apparatus to test Stage 1B: does neutron have Transition Magnetic Moment?
 - Proposed to explain neutron "bottle" lifetime MDPI Phys, 1 (2019) 271-289

 $\mathcal{H}_{int} = \begin{pmatrix} m + \mu \boldsymbol{\sigma} \cdot \boldsymbol{B} & \boldsymbol{\eta} \boldsymbol{\sigma} \cdot [\boldsymbol{B} \pm \boldsymbol{B}'] \\ \boldsymbol{\eta} \boldsymbol{\sigma} \cdot [\boldsymbol{B} \pm \boldsymbol{B}'] & m' + \mu' \boldsymbol{\sigma} \cdot \boldsymbol{B}' \end{pmatrix}$

 Seeking funding for Stage 1C at HFIR: search for n → n' assuming no mass splitting (but small magnetic field splitting)





Summary

- Searches for dark matter and baryon number violation strongly motivated! Neutrons are under-explored territory in worldwide BLV/dark matter programs
- NNBAR promises 1000x sensitivity improvement to neutron oscillations. Significant recent engineering design progress for CDR
- Cold neutron regeneration technique used at ORNL to exclude one theory that n n' oscillations caused the neutron lifetime puzzle, with more sensitive searches for this dark matter candidate recently completed
- ORNL's High Flux Isotope Reactor provides early R&D for high sensitivity program in dedicated instrument at ESS, with HIBEAM and NNBAR

I gratefully acknowledge support from the Laboratory Directed Research and Development Program of Oak Ridge National Laboratory, managed by UT-Battelle, LLC, for the U. S. Department of Energy [project number 8215] and the U.S. Department of Energy, Office of Science, Office of Nuclear Physics [DE-AC05-000R22725].



Bonus

- About 100 (mostly clickbait) stories on the experiment
- Hundreds of thousands of tweets/other social media
- 2 posted letters
- 1 piece of art
- Response from the public
 - 1/3 enthusiastic public
 - 1/3 "alternative physicists"
 - 1/3 volunteers to go through the portal

I have a Google new alert setup for **@leahjbroussard** and every time it goes off I am excited and scared all at the same time. Did she do it?? Did she do it and now we are all going to die??

@leahjbroussard have you considered what might come through the portal?... regardless, please proceed.

@leahjbroussard Not that I don't doubt you'll make excellent progress with your experiment, but I hope that you'll be taking every precaution should 'something' try to follow the particles back. Be really cautious and have failsafes to shut it down if need be.



I'm an excited hobbyist (sharing with my 6 year old son and 4 year old daughter) and wish you the best.

Hello Goodmorning, I saw your experiment recently and just wanted to let you know if you ever reach a point you want to see if you can mirror humans. I volunteer as a guinea pig.

Find us a better timeline! You're our only hope @leahjbroussard

I really encourage you and your team to watch Stranger Things before attempting to open the gate to a mirror universe. Sent from my iPhone

Thank you!



Leah Broussard