

# Exploring the extremes with NUSTAR @ FAIR

Nasser Kalantar-Nayestanaki ESRIG/University of Groningen

> On behalf of NUSTAR Coll. (Spokespers.: Zsolt Podolyak

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#### The 11 Greatest Unanswered Questions of Physics

- 1. What is dark matter?
- 2. What is dark energy?
- 3. How were the heavy elements from iron to uranium made?
- 4. Do neutrinos have mass?
- 5. Where do ultrahigh-energy particles come from?
- 6. Is a new theory of light and matter needed to explain what

happens at very high energies and temperatures?

- 8. Are protons unstable?
- 9. What is gravity?
- 10. Are there additional dimensions?
- 11. How did the universe begin?



https://www.discovermagazine.com/the-sciences/the-11-greatest-unanswered-questions-of-physics

#### Snapshot of the nuclear landscape



## Nuclear and astrophysics meet





#### Spear points of NUSTAR Phase 1

• Heavy nuclei, 3<sup>rd</sup> r-process peak

• High excitation energies



• Study of exotics



Experimental Chart of Nuclides 2975 isotones

## **NUSTAR collaboration**





Status: October 8, 2023

#### NUSTAR – The project 1.2



	Super-FRS	RIB production, separation, and identification
PSP	Experiment	Description
1.2.2	HISPEC/ DESPEC	In-beam $\gamma\text{-spectroscopy}$ at low and intermediate energy, n-decay, high-resolution $\gamma\text{-},\ \beta\text{-},\ \alpha\text{-},\ p\text{-},\ spectroscopy}$
1.2.3	MATS	In-trap mass measurements and decay studies
1.2.4	LaSpec	Laser spectroscopy
1.2.5	R <sup>3</sup> B	Kinematically complete reactions with relativistic radioactive beams
1.2.6	ILIMA	Large-scale scans of mass and lifetimes of nuclei in ground and isomeric states
1.2.10	Super-FRS	High-resolution spectrometer experiments
1.2.11	SHE	Synthesis and study of super-heavy elements
1.2.8	ELISe(*)	Elastic, inelastic, and quasi-free eA scattering
1.2.9	EXL(*)	Light-ion scattering reactions in inverse kinematics

(\*) NESR required – alternative/intermediate "operation" within MSV under discussion. SHE physics case to be evaluated.

## MSV and plans of NUSTAR



# The Physics Program



#### Overarching physics case: the creation of the (heavy) chemical elements



Big physics question requiring information on:

Equation of State Limits of existence Lifetimes, Masses P<sub>xn</sub> values Fission Reactions in star environments



#### Complementarity of NUSTAR experiments



	Super-FRS	HISPEC/DESPEC	LASPEC	MATS	R3B	ILIMA	SHE	ELISe	EXL
Masses	Precision meas.	Q-values, isomers		dressed ions,	unbound nuclei	bare ions,	precision		
				highest precision		mapping study	mass of SHEs		
Half-lives	psns-range	dressed ions,			resonance width,	bare ions,	µsdays		
		μ <b>ss</b>			decay up to 100ns	msyears			
Matter radii	interaction x-				interaction x-				matter densitiy
	section				section				distribution
Charge radii	charge-changing		mean square		charge-changing			charge density	
	cross sections		radii		cross sections			distribution	
Single-	high resolution,	high-resolution	magnetic	evolution of shell	quasi-free	evolution of	shell structure		low momentum
particle	angular	particle and γ-ray	moments,	str., pairing int.,	knockout, short-	shell closures,	of SHEs		transfers
s truc ture	momentum	spectroscopy	nucl. spins	valence nucl.	range and tensor	pairing corr.			
Collective		electromag.	quadrupole	halo structure	dipole response	changes in		electromag.	monopole
behavior		transitions	moments			deformation		transitions	resonance
EoS					polarizability,			neutron skin 🗲	neturon skin,
					neutron skin				Compressibility
Exotic	bound mesons,								
S ys tems	hypernuclei,								
	nucleon res.								

### What are the highlights of FAIR Phase 1 program?

- Understanding the 3<sup>rd</sup> r-process peak by means of comprehensive measurements of masses, lifetimes, neutron branchings, dipole strength, and level structure along the N=126 isotones;
- Equation of State (EoS) of asymmetric matter by means of measuring the dipole polarizability and neutron-skin thicknesses of tin isotopes with N larger than 82 (in combination with the results of the first highlight);
- Exotic hypernuclei with very large N/Z asymmetry.

## "PARTS" needed

#### NUSTAR experimental areas, ESSENTIAL to run!



Rich program due to approximately 2000 h beam time for NUSTAR experiments per year!

RARE-ISOTOPE BEAM FACILITIES



Exploring the extremes with NUSTAR@FAIR

#### Improvements of Radioactive Ion Beam (RIB) production with the Super-FRS at NUSTAR



# Strategy

### Definition of NUSTAR experiment phases

#### Phase 0

- R&D and experiments to be carried out with present facilities <u>and</u> FAIR/NUSTAR equipment
- Phase 1
  - Core detectors and subsystems completed
  - First measurements with FAIR/Super-FRS beams
  - Carry out experiments with highest visibility as part of the core program and within the FAIR MSV

#### • Phase 2

- FAIR evolving towards full power
- Completion of experiments within MSV
- Essentially the full program of MSV can be performed

#### • Phase 3

 Moderate projects, which have been initiated on the way (outside MSV) can be included (e.g. experiments related to return line for rings)

#### Phase 4

Major new investments and upgrades for all experiments

#### Phase 1



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#### ... to reality

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#### APPA EXPERIMENTS

NUSTAR EXPERIMENTS

September 2021 (drone video)

#### ... to reality

T

#### APPA EXPERIMENTS

NUSTAR EXPERIMENTS

HEB

September 2021 (drone video)

EB

supply building

## And in April 2024



#### NUSTAR Physics start up at FAIR



#### NUSTAR – The project 1.2



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#### Phase 1 Physics with R3B setup:

Dipole strength Distributions in heavy neutron-rich nuclei

• core vs. neutron skins & halos  $\rightarrow$  density / asymmetry



S. Bacca et al. PRL **89** (2002) 052502 PRC **69** (2004) 057001

access to EoS (e.g. neutron star) & low lying E1 strength (r-process)





### Start version for Phase 0





CALIFA (Sweden, Spain, Germany, Russia): Barrel without backward part ready



## FAIR Phase-0: <sup>12</sup>C+<sup>12</sup>C benchmark case



A benchmark case to provide very precise data that minimize the uncertainty associated to the reaction models

#### Exp. precision +-0.387%

Determined by the so-called Transmission method

$$\sigma_R = \sigma_{inel} + \sigma_I$$

$$\sigma_R = -\frac{1}{N_t} ln\left(\frac{R_i}{R_o}\right)$$

 $R_{i/o}$  is the ratio of non-interacting nuclei after target and incoming nuclei for target in (out)



L. Ponath et al., submitted to PLB

Precision of +-0.387% achieved, while the **Glauber model** including inmedium effects of Pauli blocking, Fermi motion, higher-order Eikonal corrections, Coulomb repulsion, and nuclear excitation of giant resonances **overestimates** the cross section **by around 2.5%** at higher beam energies.



Most precise constraints on the neutron pressure around saturation density from measurements of **neutron skins** and **dipole polarizabilities** 









Measured Z (red) and A (green) fission yields in actinides and pre-actinides together with the neutron-rich (dark blue circles and dots) and other nuclei (light blue circles) that can be investigated at FAIR.



## ES and FS: short range correlations



M. Duer et al. (CLAS Collaboration), Nature, 560:617, 2018

Double ratio of high-to-low momentum protons (neutrons) in nucleus A with respect to carbon are marked with dot (squares), and corresponding calculations with rectangles.



#### **Recent Achievements**



Modifications in Cave C to accomodate R<sup>3</sup>B equipment for FAIR Phase-0

- Target area (CALIFA+ LT3)
- GLAD (including Vacuum chamber)
- NeuLAND
- Tracking detectors (beampipe)



Many experiments approved by GSI-PACs have been running since 2017 with different parts installed.

## HISPEC/DESPEC

#### DESPEC



#### DESPEC (FAIR Phase-0) campaign at other labs

FATIMA fast-timing array in GANIL and Orsay, France

## AGATA+VAMOS+FATIMA

AIDA implantation and decay detector & BELEN neutron counter in RIKEN, Japan







Hybrid array for fast-timing measurements









- Evolution of the shell structure & exotic nuclear shapes in **uncharted nuclear territory**
- Spectroscopic information for the nucleosynthesis of heavy nuclei
- Comprehensive decay information at beam yields as low as one ion per hour
- Primary focus on GSI-FAIR uniqueness for nuclei around N~126, while providing competitive data on key nuclei also in other regions of the nuclear chart: around <sup>100</sup>Sn and <sup>132</sup>Sn, rare earth nuclei, ...



New Hybrid Ge-LaBr3 array in operation from April 2024



## MATS + LaSpec



Needs the low-energy cave

# Resolving isomers with ion motional phases in a Penning trap

Demonstration of isomeric cleaning with the novel phase-sensitive PI-ICR technique at JYFLTRAP. Isomerically clean beam of <sup>127</sup>Cd provided to posttrap TASISPEC decay setup.

<sup>127</sup>Cd and <sup>127m</sup>Cd: ~280 keV mass difference ( $T_{1/2}$  ~300 ms). With 250 ms excitation pattern, a maximal 180° separation achieved. With subsequent excitation, the state of interest is centered while unwanted ones are pushed further out.



Roentdek DLD40 MCP with delay-line anode

MATS

MATS





# First online application of LaSpec's new detection region



Forschungsgemeinschat

Photomultiplier Tube Scan) Adjustable Compound Apertures Parabolic Concentrator Signal Rate (Photons **Elliptical Reflector** 緣

### ILIMA

#### New masses around N=82





- High precision isochronous condition
  - Mass resolution ~10<sup>-6</sup>
- Single-ion sensitivity of new Schottky



Next exp. (98Mo and 98Zr) in May 2024

#### Isochronous mass measurements

- Neutron-rich heavy nuclei
  - Recently-developed Schottky detection in isochronous optics
  - Fast measurement with high resolution (10<sup>-6</sup>) → access to extremely short-lived nuclei (few 10ms)
- Beta-delayed neutron probabilities
  - Neutron-rich iodine isotopes (proof of principle)

#### Investigations of exotic radioactive decays

ILIMA

- Hyperfine-resolved electron capture decays
  - <sup>111</sup>Sn few electron system becomes allowed or forbidden decay
- NEEC (Nuclear excitation by electron capture)
   In cooperation with SPARC!



## Super-FRS







Secondary reactions  ${}^{70}\text{Br} \rightarrow {}^{69}\text{Br} \rightarrow {}^{68}\text{Se+}p$  ${}^{74}\text{Rb} \rightarrow {}^{73}\text{Br} \rightarrow {}^{72}\text{Kr+}p$ 

## FAIR WASA@FAIR (installed first at FRS-S2)





#### Hypernuclei:

**Combination of WASA with FRS** provides unique setup for exclusive measurements:

- FRS for high resolution spectroscopy of forward particles
- WASA for decay particles



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## WASA@FAIR (installed first at FRS-S2)







# FAIR-0: Gamov-Teller Strength at N=50 and the puzzle of <sup>100</sup>Sn mass





A. Mollaebrahimi et al., Phys.Lett. B 839, 137833 (2023)

First direct mass measurements of <sup>98</sup>Cd and <sup>97</sup>Rh with the FRS Ion Catcher

#### <sup>100</sup>Sn mass:

New results in discrepancy of <sup>100</sup>**Sn Q<sub>EC</sub> values** (Hinke et al. [1] and Lubos et al. [2]) • In recent work Mougeot et al. [3] derive the

 In recent work Mougeot et al. [3] derive the mass of <sup>100</sup>Sn from mass measurements of <sup>99-</sup>
 <sup>101</sup>In and published <sup>100</sup>Sn Q<sub>EC</sub> values
 → value of Hinke et al. is favored

#### • This work:

Evolution of shifted two-neutron shell gap at N=50:

 $\rightarrow$  Value of Hinke et al. [1] is favored.

Evolution of Gamov-Teller Strength at N=50:

 $\rightarrow$  *v*alue of Lubos et al. [2] is favored.

## Overall situation unclear, further experiments required.

[1] C.Hinke et al., Nature **486** (2012) 341
[2] D.Lubos at al., PRL **122** (2019) 222502
[3] M.Mougeot et al., Nature Phys. **17** (2021) 1099

Fire Quasi-real-time range monitoring in hadron therapy using Super-FRS EC



#### Best candidate? <sup>16</sup>O beam -> <sup>15</sup>O 43 mb 122 s <sup>14</sup>O 1.2 mb 71s

<sup>12</sup>C beam -> <sup>11</sup>C 47 mb 1221 s <sup>10</sup>C 4.3 mb 19 s

Sivaji Purushothaman et al., Sci Rep 13, 18788 (2023) Quasi-real-time range monitoring by in-beam PET: **a case for** <sup>15</sup>O

BIOMAT

(also hadron therapy with <sup>11</sup>C on mouse in Feb. 2024)

# ES and FS: new isotopes, radii and momentum distributions

#### Search for new isotopes

Longitudinal momentum distribution  $P_{\parallel}$   $\rightarrow$  Evolution of nucleon orbitals, evolution of shell structure, spectroscopic information for *r*-process

Reaction cross section  $(\sigma_R) \rightarrow$  matter radius Charge changing cross section  $(\sigma_{CC}) \rightarrow$  proton  $\rightarrow$ Identifying skins (or halos) in heavy nuclei, test of models

Neutron skin systematics → EoS of asymmetric nuclear matter

ES, FS: focus on N~82, 126 in heavy neutron-rich nuclei









ES and FS: Beta-delayed single- and multiple-neutron emission probabilities (P<sub>xn</sub>)



Important for:

r-process nucleosynthesis

nuclear structure

nuclear reactor safety

Novel method for measuring  $P_{xn}$ , simultaneously with mass,  $Q_{bxn}$ ,  $S_{xn}$  and  $T_{1/2}$ 

Complementary to worldwide programs, especially suited for <u>multi-neutron</u> emission

Enhanced transmission and separation of Super-FRS + new CSC increases P<sub>2n</sub> sensitivity by **more than 2 orders of magnitude(!)**, while reducing background by order(s) of magnitude

→ ≈30...40 P<sub>2n</sub> measurements expected during <sup>®</sup> Early Science

**FS: MNT** reaction studies with radioactive beams





Superheavy elements: "Island of Enhanced Stability"

#### See Next Talk



O. Smits, Ch.E.Düllmann et al. Nat. Rev. Phys. 6 (2024) 86

#### Message **B**:

Theory differs by 11 orders in describing  $T_{1/2}$  of known <sup>284</sup>Cn

#### Snapshot of the nuclear landscape



#### Complementarity of NUSTAR experiments



	Super-FRS	HISPEC/DESPEC	LASPEC	MATS	R3B	ILIMA	SHE	ELISe	EXL
Masses	Precision meas.	Q-values, isomers		dressed ions,	unbound nuclei	bare ions,	precision		
				highest precision		mapping study	mass of SHEs		
Half-lives	psns-range	dressed ions,			resonance width,	bare ions,	µsdays		
		μ <b>S</b> S			decay up to 100ns	msyears			
Matter radii	interaction x-				interaction x-				matter densitiy
	section				section				distribution
Charge radii	charge-changing		mean square		charge-changing			charge density	
	cross sections		radii		cross sections			distribution	
Single-	high resolution,	high-resolution	magnetic	evolution of shell	quasi-free	evolution of	shell structure		low momentum
particle	angular	particle and γ-ray	moments,	str., pairing int.,	knockout, short-	shell closures,	of SHEs		transfers
s truc ture	momentum	spectroscopy	nucl. spins	valence nucl.	range and tensor	pairing corr.			
Collective		electromag.	quadrupole	halo structure	dipole response	changes in		electromag.	monopole
behavior		transitions	moments			deformation		transitions	resonance
EoS					polarizability,			neutron skin 🗲	neturon skin,
					neutron skin				Compressibility
Exotic	bound mesons,								
S ys tems	hypernuclei,								
	nucleon res.								

# Thank you!

# Conclusions



# Funding

#### Evolution of NUSTAR project funding



### NUSTAR MSV – funding status



#### to be assigned

#### Status: March, 2021

- funding (secured and expected) from: (FAIR funding in bold face)
  - Australia
  - Belgium
  - Bulgaria
  - Canada
  - China
  - Czech Republic
  - Finland
  - France
  - Germany
  - Hungary
  - India

- Israel
- Japan
- Netherlands
- Poland
- Romania
- Russia
- Slovenia
- Spain
- Sweden
- Turkey
- United Kingdom

	NUSTAR sub-system	TDR	Cost [k€ 2005]	Funding	Construction	Date completion	Test/ Commissioning
	LEB infrastr.		2,109			07/2025	
	HISPEC/DESPEC		10,781			09/2024	
Day 1	MATS		1,173			08/2024	
	LaSpec		253			05/2024	
	R3B		17,800			07/2024	
	ILIMA		1,101			07/2025	
	Super-FRS Exp		398			12/2023	
		91.1%	33 615	93.4%	59.9%		43.9%
		value weighted	55,015	secured	value weighted		value weighted
Change since report 2021-I		0.0%	0.0	0.0%	+ 0.3%		+ 1.6%

#### Status: March 2021

## Timeline

#### NUSTAR Overall schedule: From Phase-0 to FAIR MSV



### EXL

## Recent highlight: Inelastic alpha scattering off <sup>58</sup>Ni (100 MeV/u)

#### Giant Monopole Resonance of <sup>58</sup>Ni



#### **<u>First</u>EXL**<u>pilot experiment at ESR</u> sets the world records:

- Lowest c.m. angle measured in inverse kinematics
- Most accurate extraction of monopole strength in inverse kinematics

With only one detector !!!



centroid [MeV]	EWSR [%]	
20.5(6)	79 <sup>+12</sup> -11	present data
$21.5^{+3.0}_{-0.3}$	74 <sup>+22</sup> -12	PRC <b>61</b> , 067307 (2000)
$20.8^{+0.9}_{-0.3}$	85 <sup>+13</sup>	PRC <b>73</b> , 014314 (2006)
21.1	94	RPA calculation [4]

[4] G. Colò et al, Comput. Phys. Commun. 184 (2013)

Published Oct. 2016: J.C. Zamora et al., Phys. Lett. B 763 (2016) 16

# **EXL** Phase 0 program (2018/19)

#### Giant Monopole Resonance of <sup>56</sup>Ni

Upgrade of detection system:

- Three more detectors plus new readout
- Closer geometry
- Detection system for recoil
- → Increase of solid angle substantially
- → Further reduced background

# → First measurement of the Giant Monopole Resonance in an unstable nucleus will be possible already in 2018!